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Space-time ETAS models and an improved extension

Yosihiko Ogata *, Jiancang Zhuang

The Institute of Statistical Mathematics, Minami-Azabu 4-6-7, Minato-Ku, Tokyo, 106-8569, Japan

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Abstract

For sensitive detection of anomalous seismicity such as quiescence and activation in a given region, we need a suitable statistical reference model that represents a normal seismic activity in the region. The regional occurrence rate of the earthquakes is modeled as a function of previous activity, the specific form of which is based on empirical laws in time and space such as the modified Omori formula and the Utsu–Seki scaling law of aftershock area against magnitude, respectively. This manuscript summarizes the development of the epidemic type aftershock sequence (ETAS) model and proposes an extended version of the best fitted space–time model that was suggested in Ogata [Ogata, Y., 1998. Space–time point-process models for earthquake occurrences, Ann. Inst. Statist. Math., 50: 379–402.]. This model indicates significantly better fit to seismicity in various regions in and around Japan.

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1. Introduction

Seismic quiescence and activation have been attracting attention as the precursors to a large earthquake, possibly providing useful information on its location, time and/or size (Inouye, 1965; Utsu, 1968; Ohtake et al., 1977; Wyss and Burford, 1987; Kisslinger, 1988; Keilis-Borok and Malinovskaya, 1964; Sekiya, 1976; Evison, 1977; Sykes and Jaume, 1990). Ohtake (1980) and Kanamori (1981) reviewed the studies of seismic quiescence, illustrating gaps in space–time earthquake occurrences, and hypothesizing physical mechanisms. On the other hand, Lomnitz and Nava (1983) argued that quiescence is merely due to the reduction of aftershocks of previous large earthquakes. They simulated a space–time cluster process to illustrate deceptive seis-

* Corresponding author. *E-mail address:* ogata@ism.ac.jp (Y. Ogata). mic gaps and quiescences, and claimed that these provide little predictive information about the occurrence time or the magnitude of the next large event.

Thus, it has been difficult to discuss instances of quiescence clearly in the presence of complex aftershock activity. Also, the quiescence does not always appear clearly, especially in periods and areas where the activity is high. The appearance of the quiescence also depends on a threshold magnitude of the earthquakes in the data. Indeed, the recognition of seismic anomalies appears to be subjective and still seems under development, and even controversial.

To overcome these difficulties we first need to use a practical statistical space-time model that represents the ordinary seismic activity, rather than carrying out a declustering algorithm that removes aftershock events from a catalog. Using a full homogeneous dataset, the successful model should enable us to detect anomalous temporal deviations of the actual seismicity rate from

that of the modeled occurrence rate. Indeed, the temporal ETAS (Epidemic Type Aftershock Sequence) model that ignores the spatial factor, has successfully detected quiet periods relative to the modeled rate by the transformation of occurrence times (Ogata, 1988, 1989, 1992, 2001; Ogata et al., 2003a).

There are already a number of space-time pointprocess models that take aftershock clusters into account (Kagan, 1991; Vere-Jones, 1992; Musmeci and Vere-Jones, 1992; Ogata, 1993, 1998; Rathbun, 1993, 1994; Schoenberg, 1997; Console and Murru, 2001; Zhuang et al., 2002; Console et al., 2003). In particular, the epidemic type aftershock sequence model (ETAS model; Ogata, 1988) is extended by Ogata (1998) to several space-time models that are constructed based on both empirical studies of spatial aftershock clustering with some speculative hypotheses. Goodness-of-fit of these are compared in order to judge the best practical space-time model among them for the application to both offshore interplate activity and intraplate activity inland, in and around Japan. The optimal space-time ETAS model is then extended to the hierarchical Bayesian model (Ogata et al., 2003b; Ogata, 2004). However, the diagnostic analysis based on the stochastic declustering algorithm reveals a significant bias in the spatial scaling factor in the model (Zhuang et al., 2004). The purpose of the present paper is therefore to propose a better fitted model that reduces the bias.

2. Development of the ETAS model

2.1. Epidemic type aftershock sequence model

The typical aftershock decay is represented by the Modified Omori function,

$$v(t) = K(t+c)^{-p}, (K, c, p; \text{parameters}),$$
(1)

initiated by the main shock at time origin t=0. This formula was proposed by Utsu (1957, 1961) from fits to many datasets as the extension of the Omori law (Omori, 1894). This formula remains the most widely used model for typical aftershock rate decay. To estimate the coefficients, Ogata (1983) proposed a method which maximizes the log likelihood function

$$\ln L(\theta) = \sum_{i=1}^{N} \ln v(t_i) - \int_{S}^{T} v(t) dt, \quad \theta = (K, c, p),$$
(2)

with respect to K, c and p, where $\{t_i, i=1,2,...,N\}$ is a series of occurrence times of aftershocks in the time interval (S,T) days. Typically, the Modified Omori

formula holds for quite a long period in the order of some tens of years or more, depending on the background seismicity rate in the neighboring area. See Utsu et al. (1995), and Ogata and Shimazaki (1984).

As we consider small aftershocks, however, occurrence time clustering of the events within the sequence becomes apparent. Thus, aftershock activity is not always best predicted by the single Modified Omori function, especially when it includes the conspicuous secondary aftershock activities of large aftershocks, as demonstrated in Guo and Ogata (1997) and Ogata et al. (2003a). Indeed, we see cascading complex features of aftershocks, such as interactively triggered aftershocks, including those among off-fault regions, as discussed in Felzer et al. (2002). Therefore, we assume that every aftershock can trigger further aftershocks or remote events, and that the occurrence rate at time t is given by a (weighted) superposition of the modified Omori functions shifted in time

$$\lambda_{\theta}(t) = \mu + \sum_{\left\{j: t_j \le t\right\}} e^{\alpha \left\{M_j - M_c\right\}} v(t - t_j), \tag{3}$$

where μ (shocks/day) represents the rate of the background seismicity, and the summation is taken over every *j*-th aftershock occurred before time t (days). The weighted size of its aftershocks is made as the exponential function of its magnitude M_i in accordance with the study by Utsu (1970), where M_c represents the cut-off magnitude of the fitted data. The coefficient α (magnitude⁻¹) measures the efficiency of a shock in generating its aftershock activity relative to its magnitude. For example, the α -value for Japanese earthquake swarm activity has been found to scatter in the range [0.35, 0.85], in contrast to nonswarm activity which is characterized by higher values, namely, in [1.2, 3.1] (Ogata, 1992). Note that K (shocks/day) in the v-function represents the standardized quantity by $\exp\{\alpha(M_i - M_c)\}$, which measures the productivity of the aftershock activity during a short period right after the mainshock (cf. Utsu, 1970; Reasenberg and Jones, 1994). We call Eq. (3) the ETAS (epidemic-type aftershock sequence) model, which was originally proposed to model the general seismic activity in a given region (Ogata, 1988, 1992), but which may also accurately be applied to an aftershock sequence itself (Ogata, 1989, 2001; Guo and Ogata, 1997; Ogata et al., 2003a).

For a sequence of occurrence times with associated magnitudes, we can estimate the parameters $\theta = (\mu, K, c, \alpha, p)$ of the ETAS model that are common to all *i*, by maximizing the log-likelihood function that is of the

same form as the one in Eq. (2) with the exception that the Modified Omori intensity function v(t) is replaced by the occurrence rate $\lambda_{\theta}(t)$ of ETAS model. See Utsu and Ogata (1997) for computational codes and technical aspects, and Helmstetter and Sornette (2002), for example, for some discussions of the statistical features of the ETAS model.

2.2. Space-time ETAS model

Consider point-process models for the data of occurrence times and locations of earthquakes, whose magnitudes are equal to or larger than a certain threshold (cut-off) magnitude M_c . The occurrence rate $\lambda(t, x, y|H_t)$ of a space-time point process is mathematically defined in terms of the occurrence probability of an event at time t and the location (x, y) conditional on the past history of the occurrences such that

Prob{an event in
$$[t, t + dt] \times [x, x + dx] \times [y, y + dy]|H_t$$

= $\lambda(t, x, y|H_t)dtdxdy + o(dtdxdy)$

where $H_t = \{(t_i, x_i, y_i, M_i, S_i); t_i \le t\}$ is the history of occurrence times $\{t_i\}$ up to time t, with corresponding epicenters $\{(x_i, y_i)\}$, magnitudes $\{M_i\}$ and 2×2 -matrices $\{S_i\}$ for anisotropic clusters (see the model Eqs. (5)~(7) below and Appendix). Hereafter, for simplicity, we write $\lambda(t, x, y)$ in place of $\lambda(t, x, y|H_t)$. Also, the space-time ETAS model is an extension of the above ETAS model, satisfying the relation $\int \int \lambda(t,x,y) dx dy =$ $\lambda(t)$ where $\lambda(t)$ represents the ETAS model (3). Consider a space-time occurrence intensity $\lambda_{\theta}(t, x, y)$ at time and location (t, x, y), where $\lambda_{\theta}(t, x, y) dt dx dy$ is the probability of an earthquake occurring in the infinitesimal space-time volume $[t, t+dt) \times [x, x+dx) \times [y, t+dt]$ y+dy). As in the typical space-time extensions of the ETAS model, Ogata (1998) considered the separable form in time and space clustering effects, and compared the following models expressed by

$$\lambda_{\theta}(t, x, y) = \mu(x, y) + \sum_{\substack{\{j: t_j \leq t\}}} v(t - t_j)$$
$$\times g(x - x_j, y - y_j; M_j - M_c), \tag{4}$$

where

$$g(x - x_{j}, y - y_{j}; M_{j} - M_{c}) = \exp\left[-\frac{1}{2} \frac{(x - x_{j}, y - y_{j})\mathbf{S}_{j}(x - x_{j}, y - y_{j})^{t}}{de^{\alpha(M_{j} - M_{c})}}\right]$$
(5)

or

$$g(x - x_{j}, y - y_{j}; M_{j} - M_{c}) = \frac{e^{\alpha(M_{j} - M_{c})}}{\left[(x - x_{j}, y - y_{j}) \mathbf{S}_{j} (x - x_{j}, y - y_{j})^{t} + d \right]^{q}}$$
(6)

or

$$g(x - x_j, y - y_j; M_j - M_c) = \left[\frac{(x - x_j, y - y_j)\mathbf{S}_j(x - x_j, y - y_j)^t}{e^{\alpha(M_j - M_c)}} + d\right]^{-q}$$
(7)

where S_i is an adimensional 2×2 positive definite symmetric matrix, and $(x, y)^t$ indicates the vector transpose. The square brackets have dimension of degree² where "degree" corresponds to the global distance in latitude (i.e., 111.11 km) throughout the present paper. The quadratic form within the brackets indicates that the aftershocks are spatially distributed with ellipsoidal contours. Indeed, Utsu and Seki (1955) and Utsu (1969) used an ellipsoid to measure aftershock area instead of a rectangle. Such an ellipsoid reflects not only the approximate shape of the ruptured fault and its dip angle, but also the location errors of the aftershock hypocenters. The logarithm of the aftershock area is highly correlated with the main shock's magnitude, leading to the famous Utsu-Seki law, the extension of which is represented by the denominator $e^{\alpha(M_i-M_c)}$ in the model (4). The inverse power decay of the aftershock distribution against the distance takes the remote triggering phenomena into account.

It is often the case that the epicenter of a mainshock is located on the margin of its aftershock area, because the epicenter corresponds to the location of earthquake fracture initiation. For such a mainshock, the epicenter location in the catalog is not suitable in the model in (4). Therefore, we replace the mainshock's epicenter coordinates with the centroid of the aftershocks (the mean coordinates of the aftershocks) for the model (4). Such centroid of aftershocks could be closely related to the centroid of the ruptured fault determined in the Harvard CMT catalog due to Dziewonski et al. (1981). Also, spatial distributions of aftershock epicenters are not usually isotropic owing to the aforementioned reasons. The compiling procedure for the centroid of clusters and also the matrices S_i representing the ellipsoid of anisotropic clusters are summarized in the Appendix. Such a recompiled dataset is useful in the model (4) for some significantly large earthquakes, whereas for the remaining supermajority of events, the

centroid can be taken as the epicenter of the original catalog, and S_i as the identity matrix.

Given the recompiled dataset of origin times and space coordinates of earthquakes together with their magnitudes and matrices $\{(t_i, x_i, y_i, M_i, \mathbf{S}_i); M_i \ge M_c, i=1,...,n\}$ during a period [0, *T*] and in region A, we can calculate the log-likelihood function of the parameter $\theta = (\mu, K, c, \alpha, p, d, q)$ characterizing the space–time point-process model, which is given by

$$\ln L(\theta) = \sum_{i=1}^{N} \ln \lambda_{\theta}(t_i, x_i, y_i) - \int_{S}^{T} \iint_{A} \lambda_{\theta}(t, x, y) dt dx dy.$$
(8)

Interested readers are referred to Daley and Vere-Jones (2002; Section 7) for the derivation of this formula and Ogata (1998) for the numerical calculations. The maximum likelihood estimate (MLE) $\hat{\theta} =$ $(\hat{\mu}\hat{K}, \hat{c}, \hat{\alpha}, \hat{p}, \hat{d}, \hat{q})$ is the one that maximizes the function. The physical dimensions of these parameters are provided in Table 1.

For the comparison of goodness-of-fit of the competing models to a dataset, Akaike's Information Criterion (AIC; Akaike, 1974) is useful. The statistic $AIC = -2 \ln L(\theta) + 2\dim(\theta)$ is computed for each of the models fitted to the data. In comparing models with a different number of parameters, adding the quantity $2\dim(\theta)$ roughly compensates for the additional flexibility which the extra parameters provide. The model with the lower AIC-value is taken as giving the better choice for forward prediction purposes. Insofar as it depends on the likelihood ratio, the AIC can also be used as a rough guide to the model testing. As a rule of thumb, in testing a model with k+d parameters against a null hypothesis with just k parameters, we take a difference of 2 in AIC values as a rough estimate of significance at the 5% level.

Ogata (1998) compared the goodness-of-fit of the three cases (5)–(7) of the models for the following cases where:

- C1. Homogeneous Poisson field for the background seismicity, $\mu(x,y)=\mu=\text{const.}$, and *isotropic* clustering, i.e., $\mathbf{S}_i=2\times 2$ *identity matrix*;
- C2. Non-homogeneous Poisson field for the background seismicity, $\mu(x,y) = v\mu_0(x,y)$ where $\mu_0(x,y)$ is a baseline-spline-surface, and *isotropic* clustering, i.e., $S_j = 2 \times 2$ *identity matrix*;
- C3. Homogeneous Poisson field for the background seismicity, $\mu(x,y) = v\mu_0(x,y)$ where $\mu_0(x,y)$ is a baseline-spline-surface, and anisotropic clustering, $\mathbf{S}_j = 2 \times 2$ *positive-definite symmetric matrix* depending on *j* (cf., Appendix) and the center of each cluster (x_j, y_j) is modified to be the centroid (average) of the coordinates of the aftershocks.

According to Ogata (1998) the AIC always selected the model (4) with (7), the consequences of which are as follows:

R1. The triggered clusters in space extend beyond the traditional aftershock regions, indicating a much

Table 1

The MLEs of space-time ETAS model fitted to the three datasets 1926-1995

Model	$\hat{\mu}$	Ŕ	ĉ	â	Ŷ	p̂	â	\hat{q}	AIC
Unit	Events/da	Events/day/degree ²		Magnitude ⁻¹			Degrees ²		
Off the e	ast coast of Tohoku	District (Region A	l) M4≥0.5, 4333 e	vents					
(7)	0.707×10^{-3}	0.967×10^{-4}	0.836×10^{-2}	1.281	1.281	0.909	0.184×10^{-2}	1.565	0.0
(10)	0.711×10^{-3}	0.988×10^{-4}	0.840×10^{-2}	1.962	1.326	0.910	0.203×10^{-2}	1.570	0.6
(11)	0.708×10^{-3}	0.106×10^{-3}	0.842×10^{-1}	1.273	(1.15)	0.910	0.222×10^{-2}	1.571	-0.6*
Central d	and western Honshi	u (Region B) $M \ge 4$.0, 3007 events						
(7)	0.335×10^{-3}	0.167×10^{-3}	0.514×10^{-2}	0.910	0.910	0.961	0.341×10^{-3}	1.405	0.0
(10)	0.337×10^{-3}	0.171×10^{-4}	0.520×10^{-2}	0.935	0.740	0.961	0.403×10^{-3}	1.408	-3.2*
(11)	0.330×10^{-3}	0.148×10^{-3}	0.512×10^{-2}	0.972	(1.15)	0.960	0.261×10^{-3}	1.397	19.0
All Japar	n data (Region C) I	M≥5.0, 4586 event	ts						
(7)	0.985×10^{-5}	0.274×10^{-3}	0.740×10^{-2}	1.113	1.113	0.910	0.434×10^{-2}	1.513	0.0
(10)	0.101×10^{-4}	0.290×10^{-3}	0.748×10^{-2}	1.154	0.891	0.911	0.552×10^{-2}	1.524	-13.2*
(11)	0.101×10^{-4}	0.247×10^{-3}	0.745×10^{-2}	1.626	(1.15)	0.910	0.413×10^{-2}	1.515	-0.7

For each dataset the AIC-values are given relative to the value of the model (7) for easier comparison. The original AIC values of the model (7) are AIC_A=32897.4, AIC_B=15836.7 and AIC_C=45161.4 for the dataset from regions A, B, and C, respectively, which are the same values given in Ogata (1998). * Indicates the smallest AIC value among the models (7), (10) and (11).

more diffuse boundary with power law decay rather than a more clearly defined region with a fairly sharp boundary converging faster than the exponential decay.

- R2. There may perhaps be two components (near field and far field) with different characteristics; the near field component corresponds to the traditional aftershock area around the ruptured fault, and the far field component may relate to the so called the 'aftershocks in wide sense' such as immigrations of earthquake activity or causal relations between distant regions, caused by tectonic changes of the stress-field due to the rupture or dynamic triggering of the seismic waves.
- R3. The cluster regions scale with magnitudes in close agreement with the Utsu–Seki formula.

3. Extension of the best fitted space-time model

Using the models (4) with (5) \sim (7), Zhuang et al. (2004) implemented the stochastic declustering of a data set from Hypocenter Catalog of the Japan Meteorological Agency (JMA) in order to make diagnostic analysis of space-time features of clusters using the space-time ETAS models. The majority of the diagnostic results show that the functions for each component in the formulation of the above best space-time ETAS model (4) with (7) confirms the superiority to, and shows graphically better fit to various declustered statistics than the model (4) with (5). In particular, it is shown that the scale of the triggering region is still an exponential law as formulated in (7). However, one of the important diagnostic features is that some systematic deviation (bias in slope) is seen from the expected number of offspring (cf. Fig. 2a in the present paper) for the considered data from a central Japan region that is defined in Section 4. This suggests us that some modification of the model is desirable.

In Ogata (1998), the common standard form including the models (5)~(7) is given by the multiplication of normalized time and space density distributions besides the multiplication of size function in such a way that

$$g(x, y; M) = \kappa(M) \times \frac{(p-1)c^{p-1}}{(t+c)^p} \times \left[\frac{1}{\pi\sigma(M)}h\left\{\frac{(x, y)\mathbf{S}(x, y)^t}{\sigma(M)}\right\}\right]$$
(9)

where h(x, y) represents the different forms in (5)~(7), and the size function is assumed to be $\kappa(M) = \text{const.} \times \sigma(M) \propto e^{\alpha M}$ in the paper. Now, the point of the present extension is to remove the constraint between $\kappa(M)$ and $\sigma(M)$. This leads to the new model

$$g(x - x_j, y - y_j; M_j - M_c) = e^{(\alpha - \gamma)(M_j - M_c)} \\ \times \left[\frac{(x - x_j, y - y_j) \mathbf{S}_j (x - x_j, y - y_j)^t}{e^{\gamma(M_j - M_c)}} + d \right]^{-q},$$
(10)

which requires the eight parameters $\theta = (\mu, K, c, \alpha, \gamma, p, d, q)$. In the following sections, we will compare this with the model (4) with (7). Furthermore, by fitting least squares to the diagnostic plots, Zhuang et al. (2004) implies that

$$\tilde{\gamma} = 0.50 \log_e 10 \approx 1.15,\tag{11}$$

rather than the MLE, $\hat{\gamma} = 1.334$, in the model (4) with (7) for the central Japan data from the JMA earthquake catalog. Actually this agrees with the famous empirical formulae $\log_{10}A = M + 4.0$ in Utsu and Seki (1955), or equivalently $\log_{10}L = 0.5 M - 1.8$ in Utsu (1961), where A and L represents the area and length of the aftershock zone, respectively, for the mainshock magnitude M. Relevantly, the above intersect constants for the land (intraplate) events are smaller than those for the sea (interplate) events, but the slopes 1.0 and 0.5 remain the same according to Utsu (1969). Related studies on the scaling relations are Shimazaki (1986), Yamanaka and Shimazaki (1990), and Scholz (1990, Section 4.3.2). Thus, it is also worthwhile to compare the version of the model (4) with (10) restricted by the fixed value in (11).

4. Application to the data sets

We use the hypocenter data compiled by the Japan Meteorological Agency (JMA), and consider three data sets from areas of tectonically distinctive features and also their mixture. Initially, data of earthquakes of magnitude (M) 4.5 and larger are chosen from the wide region $36{\sim}42^{\circ}$ N and $141{\sim}145^{\circ}$ E (Off the east coast of Tohoku District; see Fig. 1) for all depths down to 100 km and for the time span 1926-1995. From now on we refer to this region as Region A. In Region A, most of the large earthquakes took place on the plate boundary between the North American and the subducting Pacific plates. We ignore the depth axis and consider only two-dimensional locations (longitude and latitude) of earthquakes, restricting ourselves to the shallow events. Most of the events are distributed within depths down to 60 km. Another area of interest is the



Fig. 1. Epicenter of earthquakes of magnitude 4.0 and larger (depth \leq 100 km) in and around Japan, for the period 1926–1995, and regions (A), (B), (C), and (D) from which the space-time models are applied.

central and western part of Honshu Island, Japan, the region $34 \sim 38^{\circ}$ N and $131 \sim 137^{\circ}$ E (hereinafter referred to as Region B) shown in the Fig. 1, where most earthquakes are considered to be intraplate events that occur within the Eurasian plates. Shallow earthquakes ($h \le 45$ km) of M4.0 and larger are considered for the time span 1926~1995. Finally, we consider the data set of the hypocenter locations of 4586 earthquakes of M5.0 or larger in the region $30 \sim 47^{\circ}$ N and $128 \sim 149^{\circ}$ E with depths shallower than 65 km for the period from 1926 through 1995, in and around Japan, referred to in Fig. 1 as Region C.

The maximum likelihood estimations and AIC comparisons are made for every data set. In calculating the distance between the earthquake epicenters, the distance in longitude is reduced to $\cos(\pi y_0/180^\circ)$ times as large as its latitude (1° corresponds to about 111.11 km), where y_0 is taken to be the latitude of the center of the area. Table 1 lists the estimated parameters and AICs of the respective models, when the background intensity is invariant in space such that $\mu(x,y) = \mu$ in (4). The AIC comparison among the considered models shows that the goodness-of-fit of the models (4) with (7) and (10) exhibit similar performance in both areas A and B, but indicate better performance as regards to (10) for Region C. However, as discussed in Ogata (1998), the estimate $\hat{p} < 1$ for a long period indicates that the assumption $\mu(x,y)=\mu$ in (4) is inappropriate, which is indeed, common sense among seismologists.

More realistic versions of the three models are fitted to each data set. Table 2 summarizes the results for the three models with the location-dependent background intensity (4) with $\mu(x, y) = v\mu_0(x, y)$ where the baseline function $\mu_0(x, y)$ is the same as that represented by the bi-cubic spline surfaces for the corresponding regions in Ogata (1998). Moreover, we assume that the matrix S_i is the 2 × 2 identity one for any event *i*, indicating isotropic spatial clustering. Furthermore, Table 3 summarizes the results for the three models with the locabackground intensity tion-dependent and with anisotropic spatial clustering, where the matrix S_i for any event j is obtained by the procedure described in the Appendix.

The goodness-of-fit of the extended model (4) with (10) becomes significantly better than the model with (7) for all the data sets in both Tables 2 and 3. It is noteworthy that each AIC in Tables 2 and 3 is remarkably smaller than the corresponding AIC in Table 1. However, the justification is not established for a straightforward comparison of AIC of the models with the adjusted function $\hat{\mu}_0(x, y)$ beforehand, nor as regards to additional explanatory data for the anisotropic clustering. In fact, the number of coefficients of the B-spline function is 96, 198 and 480 for the background

Table 2 The MLEs of space-time ETAS model fitted to the three datasets 1926–1995

Model	ŵ	Ŕ	ĉ	ά	Ŷ	p̂	â	\hat{q}	AIC
Unit	Events/day/degree ²		Days	Magnitude ⁻¹			Degrees ²		
Off the e	ast coast of Tohoki	u District (Region A	4) M≥4.5, 4333 e	vents					
(7)	0.131×10^{-3}	0.416×10^{-4}	0.230×10^{-1}	1.605	1.605	1.043	0.103×10^{-2}	1.857	-1037.7
(10)	0.134×10^{-3}	0.402×10^{-4}	0.243×10^{-1}	1.645	1.331	1.050	0.179×10^{-2}	1.648	-1057.6*
(11)	0.134×10^{-3}	0.521×10^{-4}	0.247×10^{-1}	1.607	(1.15)	1.053	0.242×10^{-2}	1.648	-1053.3
Central d	and western Honsh	u (Region B) $M \ge 4$	4.0, 3007 events						
(7)	0.495×10^{-4}	0.511×10^{-4}	0.855×10^{-2}	1.040	1.040	1.027	0.316×10^{-3}	1.568	-662.1
(10)	0.953×10^{-4}	0.524×10^{-4}	0.878×10^{-2}	1.103	0.802	1.028	0.416×10^{-3}	1.580	-675.1*
(11)	0.954×10^{-4}	0.439×10^{-4}	0.876×10^{-2}	1.128	(1.15)	1.027	0.269×10^{-3}	1.564	-653.8
All Japa	n data (Region C) .	M≥5.0, 4586 even	nts						
(7)	0.131×10^{-4}	0.475×10^{-4}	0.176×10^{-1}	1.518	1.518	1.021	0.232×10^{-2}	1.738	-1426.7
(10)	0.134×10^{-3}	0.468×10^{-4}	0.186×10^{-1}	1.644	1.183	1.026	0.394×10^{-2}	1.800	-1435.0
(11)	0.134×10^{-3}	0.492×10^{-4}	0.186×10^{-1}	1.637	(1.15)	1.026	0.412×10^{-2}	1.799	-1436.7*

For each dataset the AIC-values are given relative to the value of the model (7) in Table 1. * Indicates the smallest AIC value among the models (7), (10) and (11).

seismicity $\mu_0(x, y)$ of Region A, B and C, respectively, and the AIC differences confirm a significantly better fit than they do with regards the homogeneous background seismicity.

The *p*-values which are smaller than 1.0 in Table 1 now become larger than 1.0, which agrees with our experience in estimating the ETAS model to various data during a long period of seismic activity. Therefore, along with the substantial decrease of each AIC in Tables 2 and 3 compared with the AIC of the corresponding models, we believe that the inclusion of the non-homogeneous background seismicity in the space-time modeling provides significantly better

performance as regards to the present three data sets. The location-dependent background intensity model also provided a larger estimated value of the coefficient q of the decay power of distance to the cluster members.

Comparing Table 2 with 3, we see that the parameter values of the corresponding models are quite similar and that the decrease of the AIC is not very large, in spite of the implicitly used parameters for the anisotropic clustering. Therefore, it is not very clear whether or not the anisotropic modeling significantly improves the goodness-of-fit, but the improvement seems to become clearer as the amount of data increases, or the

Table 3 The MLEs of space-time ETAS model fitted to the three datasets 1926–1995

Model	ŷ	Ŕ	ĉ	â	Ŷ	\hat{p}	â	\hat{q}	AIC
Unit	Events/day/degree ²		Days	Magnitude ⁻¹			Degrees ²		
Off the ed	ast coast of Tohoki	u District (Region 2	4) M≥4.5, 4333 e	events					
(7)	0.131×10^{-3}	0.382×10^{-4}	0.231×10^{-1}	1.612	1.612	1.043	0.102×10^{-2}	1.600	-1059.3
(10)	0.134×10^{-3}	0.375×10^{-4}	0.245×10^{-1}	1.657	1.326	1.050	0.182×10^{-2}	1.662	-1081.4*
(11)	0.135×10^{-3}	0.491×10^{-4}	0.248×10^{-1}	1.617	(1.15)	1.043	0.102×10^{-2}	1.600	-1077.6
Central a	und western Honsh	u (Region B) $M \ge -$	4.0, 3007 events						
(7)	0.946×10^{-4}	0.477×10^{-4}	0.866×10^{-2}	1.057	1.057	1.027	0.316×10^{-3}	1.577	-678.6
(10)	0.954×10^{-4}	0.489×10^{-4}	0.895×10^{-2}	1.126	0.804	1.029	0.423×10^{-3}	1.589	-693.8*
(11)	0.955×10^{-4}	0.406×10^{-4}	0.890×10^{-2}	1.148	(1.15)	1.028	0.274×10^{-3}	1.575	-672.3
All Japar	n data (Region C)	M≥5.0, 4586 ever	its						
(7)	0.131×10^{-3}	0.412×10^{-4}	0.176×10^{-1}	1.549	1.549	1.020	0.221×10^{-2}	1.752	-1453.6
(10)	0.135×10^{-3}	0.398×10^{-4}	0.186×10^{-1}	1.484	1.196	1.026	0.396×10^{-2}	1.828	-1521.1
(11)	0.135×10^{-3}	$0.428 imes 10^{-4}$	0.186×10^{-1}	1.673	(1.15)	1.026	0.423×10^{-2}	1.826	-1522.5*

For each dataset the AIC-values are given relative to the value of the model (7) in Table 1. * Indicates the smallest AIC value among the models), (7), (10) and (11).

threshold magnitude lowers. Finally, throughout Tables 1, 2, and 3, it is confirmed that model (4) with (10) for the spatial clustering improves the goodness-of-fit and its degree of significance depends on the region of different seismicity patterns. Furthermore, reducing the number of parameters may be possible by fixing the parameter γ in (11) depending on the data, rather than considering the restriction $\alpha = \gamma$ in the model (10), which is nothing but the model (7).

5. Diagnostic analysis by stochastic declustering

5.1. Stochastic declustering

To obtain an objectively declustered catalogue, Zhuang et al. (2002, 2004) proposed the stochastic declustering method as an alternative to the conventional declustering methods. In this method, it is no longer determined whether an earthquake is a background event or is triggered by another. Instead, each event has a probability of either being a background event or a direct offspring triggered by another. The main task of the stochastic declustering algorithm is to estimate this probability for each event according to some models used to describe earthquake clustering features.

The technical key point of the stochastic declustering method is the thinning operation to a point process (Ogata, 1981, 1998; Daley and Vere-Jones, 2002; Zhuang et al., 2004). Observing the model (4), the relative contribution of the previous *i*-th event to the total conditional intensity at the occurrence time and location of the *j*-th event, (t_i, x_i, y_i) , is

$$\rho_{i,j} = v(t_j - t_i)g(x_j - x_i, y_j - y_i; M_i - M_c)/\lambda(t_j, x_j, y_j)$$

by the definition. That is to say, for each i=1,2,...,N, select the *j*-th event with probability $\rho_{i,j}$, then we can realize a subprocess of the events triggered by the *i*-th event. In this way, $\rho_{i,j}$ can be regarded as the probability of the *j*-th event being triggered by the *i*-th event. Furthermore, the probability of the event is defined by the event *j* being a background event is $\phi_j = \mu(t_j, x_j, y_j) / \lambda(t_j, x_j, y_j)$ and the probability that the *j*-th event is triggered is $\rho_j = 1 - \phi_j = \sum_{\{i: i \le j\}} \rho_{i,j}$. In other words, if we select each event *j* with probabilities ϕ_j , we can then form up a new process, the background subprocess. For the algorithm of the stochastic declustering, the readers are referred to Zhuang et al. (2002, 2004).

A stochastically declustered catalogue produced from the above procedures is not unique, since it depends on random numbers used in the selection of events to form the background seismicity. Unlike conventional declustering methods, the stochastic declustering method does not fix judgment on whether an event is an aftershock or not. Instead, it gives a probability of how each event might look an aftershock. Namely, the stochastic declustering is understood to be a simulation, or to be precise, a bootstrap resampling, and we understand this to be advantageous because it shows the uncertainty about earthquake activity. Thus, simply by using random simulation of the thinning method, we can easily produce stochastic copies of the declustered catalogue. Stochastic declustering realizes (simulates) many possible configurations of background events depending on the seeds of random number. In this way, we can make use of graphical statistics, based on the repeated thinning realizations to discuss uncertainty and significance of interesting phenomena, in much the same way as the conventional bootstrapping does.

5.2. Distribution of distance to offspring events relative to ancestor's magnitude

In order to examine the appropriateness of the function form $\exp\{\alpha(M-M_c)\}/d$ in the model (7), Zhuang et al. (2004) calculated the distance $r_{i,j}$ between a triggered event j and its direct ancestor, event i, that belong to a given magnitude band $M_i \in \Delta M$, using the set of clusters that are declustered from the coordinates data of 8283 target events of M4.2 or larger, from the rectangular region 130–146° E and 33–42.5° N (see Region D in Fig. 1), in the depth range (0, 100) km, during the period from 10,001-th day from 1926, up until the end of 1999. Then, to estimate the scaling parameter d in (7) for each magnitude band ΔM , instead of maximizing the log-likelihood function of the possibly quite sizeable resampled data due to the stochastic declustering, Zhuang et al. (2004) consider maximizing the log-pseudo-likelihood

$$\log L(D) = \sum_{\{i; M_i \in \Delta M\}} \sum_{\{j; i < j\}} \rho_{i,j} \log \left\{ \frac{2(q-1)D^{q-1}r_{i,j}}{\left(r_{i,j}^2 + D\right)^q} \right\}$$

Fig. 2a shows the plot of \hat{D} against the magnitude band $\Delta M = (M - 0.05, M + 0.05)$ for the model (7) that is the similar to Fig. 14 in Zhuang et al. (2004). The plots should correspond to $\hat{d}e^{\hat{\alpha}(M-M_c)}$ where \hat{d} and $\hat{\alpha}$ are



Fig. 2. Circles indicate the mode of distance-distribution (i.e., D; the most frequently appeared distance from the parent to the cluster members, cf. Section 5.2) against the corresponding magnitudes of triggering (parent) earthquakes, calculated based on the stochastically declustered clusters that are reconstructed from the JMA data from Region D: the panels (a) and (c) show those where the stochastic-declustering algorithm is performed using the model (7) and (10), respectively; and the panels (b) and (d) are the same plots for a simulated data using the model (7) and (10), respectively. The solid straight lines in (a, b) and the dotted lines in (c, d) indicate the function $\hat{d}e^{\hat{z}(M-M_c)}$ of magnitude *M*, with the maximum likelihood estimate of the corresponding model, respectively; the straight lines in (c) and (d) indicate the function $\hat{d}e^{\hat{r}(M-M_c)}$.

the MLE of the model (4) with (7). However, the \hat{D} plot alignment has a significantly biased smaller slope than that of the log-plot of $\hat{d}e^{\alpha(M-M_{c})}$ as seen in Fig. 2a. On the other hand, Fig. 2c shows a similar plot obtained by the applying the presently proposed model (10), which has little systematic deviation. The significance and stability of the plots are demonstrated in Fig. 2b and d by the corresponding simulated dataset of the respective model.

6. Concluding remarks

The detected rate of earthquakes in a catalog generally changes not only with location, but also with time, due to the configuration of seismometers and changing observational environments in space and time. Therefore, we have studied the seismicity using data from the lowest threshold magnitude, above which earthquakes are completely detected, satisfying Gutenberg–Richter's law. The estimated parameter values themselves depend, by scale difference, on the magnitude thresholds of complete detection, except for the α , γ , p and q-values, in principle. However, the proportion of earthquakes smaller than the minimum threshold magnitude for the complete data is usually substantial because the number of earthquakes increases exponentially (Gutenberg–Richter's law) assuming all earthquakes are completely detected. From the viewpoint of effective use of data, this is quite wasteful in the statistical analysis of seismic activity. Thus, our next step for the practical space–time seismicity analysis is to develop the presently improved model, taking account of the space–time detection rate as a function of magnitude, time and location, as implemented in Ogata and Katsura (1993) for space–time changes of *b*-value in the Gutenberg–Richter magnitude–frequency law.

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Appendix A. Data processing for anisotropic clusters

We have proposed the model (4) with (10) as the extended version of the model (4) with (7) that was best fitted in Ogata (1998). This model shows significantly better fit to seismicity in a wide region. The difference of the goodness-of-fit depends on the region of analysis. Thus, we calculate the maximum likelihood estimates (MLE) of the space–time ETAS model, which show regional characteristics of seismicity such as the background seismicity, aftershock population sizes, aftershock productivity, aftershock decay rate, scaling of spatial clustering, etc., throughout the considered period.

In order to replace the epicenter coordinate (x_j, y_j) in the catalog by the centroid of aftershock locations for the coordinate (\bar{x}_j, \bar{y}_j) in the model (4), we estimate retrospectively using the aftershock distribution as follows. First, we identify clusters of aftershocks, $i=1,2,\ldots,N_j$, by the algorithm that is provided below. Then, we take the average (\bar{x}, \bar{y}) of the locations of the cluster members to replace the epicenter of the main shock (x_j, y_j) only when the difference is significant as determined by the statistical procedure (Ogata, 1998).

Also, the anisotropic spatial aftershock distribution represented by the matrix \mathbf{S}_j in (4) for an ellipsoid is estimated as follows. We fit a bi-variate normal distribution to the location coordinates of the aftershocks in each cluster (see below) to obtain the maximum likelihood estimate of the average vector ($\hat{\mu}_1$, $\hat{\mu}_2$) and the covariance matrix with the adimensional elements ($\hat{\sigma}_1$, $\hat{\sigma}_2$) and ρ for \mathbf{S}_j in (4) in the form

$$\mathbf{S} = \begin{pmatrix} \sigma_1^2 & \rho \sigma_1 \sigma_2 \\ \rho \sigma_1 \sigma_2 & \sigma_2^2 \end{pmatrix}.$$

This is listed in the recompiled catalog only when they are significantly different from the identity matrix as the null hypothesis (i.e., $\sigma_1 = \sigma_2 = 1$ and $\rho = 0$). Specifically, according to Ogata (1998), the minimum AIC procedure (Akaike, 1974), instead of the likelihood ratio test, is adopted among all the nested models including the null hypothesis. For the rest of events in the cluster, the null hypothesis is always adopted; namely, the same coordinate as the epicenter given in the catalog and the identity matrix for **S**_j.

The algorithm for identifying the aftershock clusters starts by selecting the largest shock in the original catalog for the mainshock. If there are multiple largest shocks with equal magnitude, the earliest one is adopted as the main shock. Then, to form a cluster, we set a space-time window with the bounds of distance and time depending on the magnitude of the main shock, based on the empirical laws of aftershocks (c.f., Utsu, 1969). For example, the algorithm in Ogata et al. (1995), except for foreshocks, describes its explicit form. The identification of the aftershock cluster is surely subjective to some degree, in spite of the method based on the empirical laws. Nevertheless, this is useful for our eventual objective to estimate the centroidal coordinates and coefficients for the anisotropy of the clusters for some large earthquakes. Indeed, we can apply a simple similar compiling procedure based on the aftershocks during only a few days' period. Also, the stochastic declustering algorithm will be useful.

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